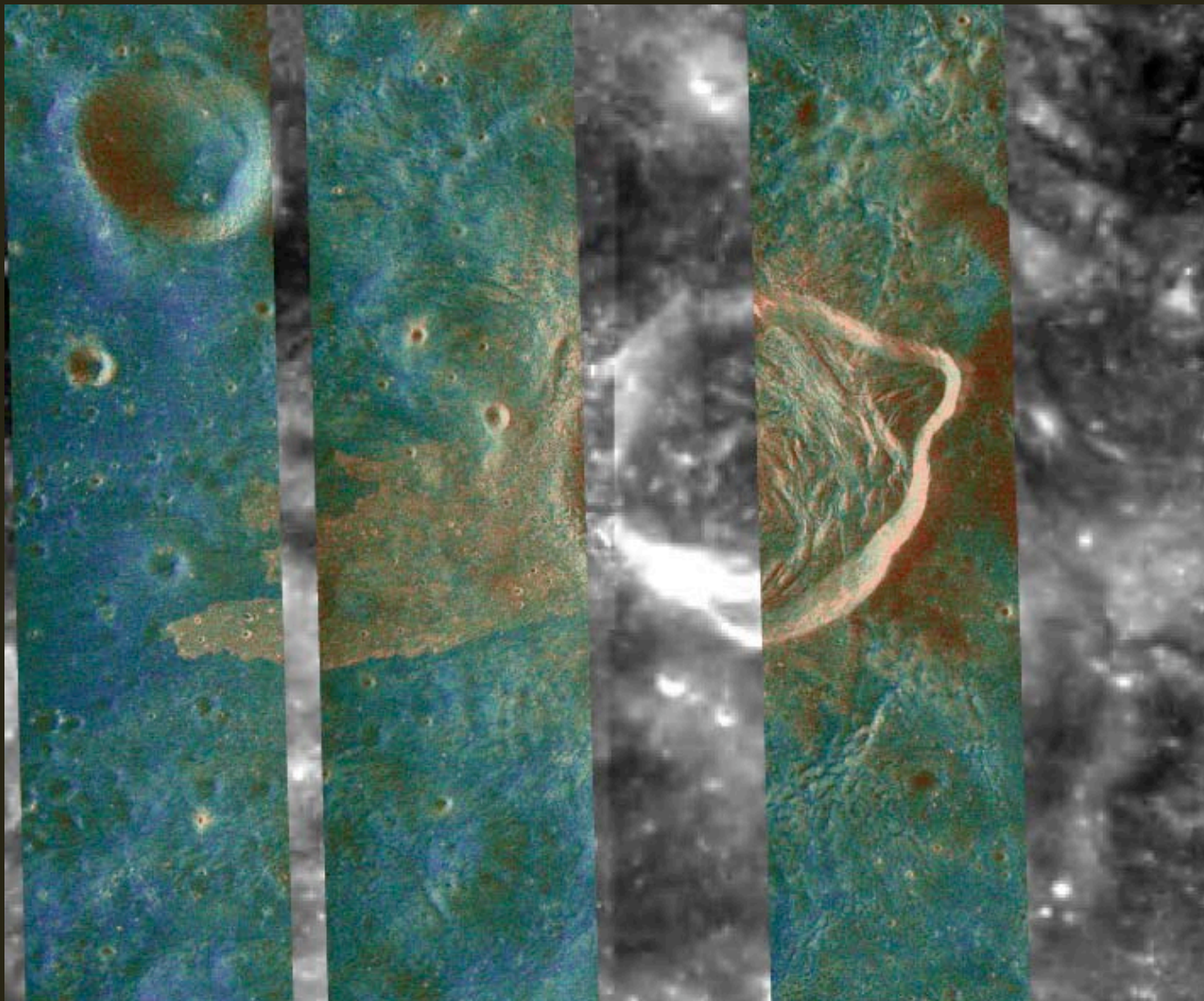


Scientific Highlights from Mini-RF

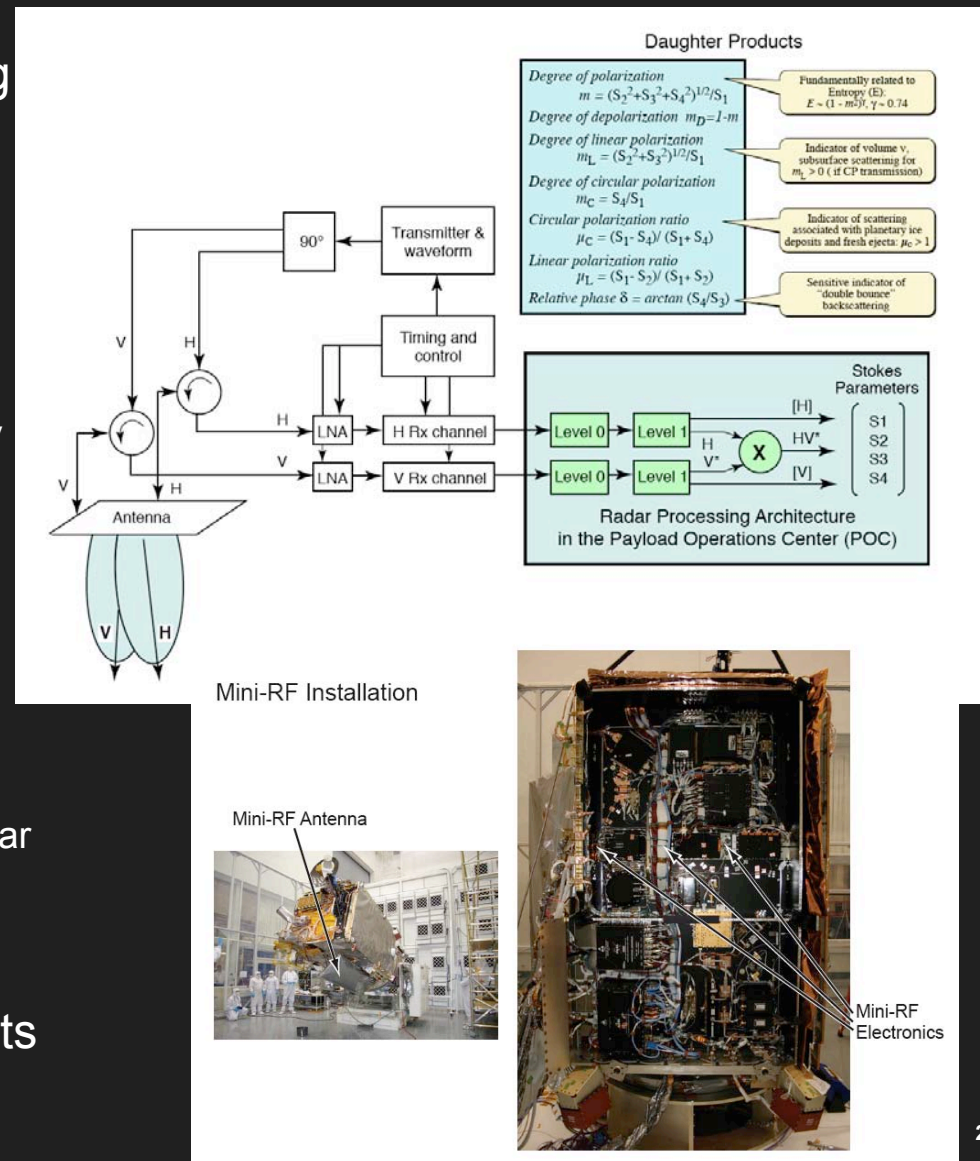


Ben Bussey
& the Mini-RF
Team

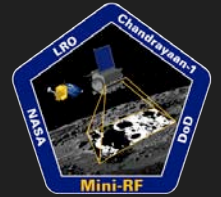
Mini-RF Instrument Description



- Two-band, two-resolution imaging radar
 - S-band ($\lambda=12.6$ cm)
 - X-band ($\lambda=4.2$ cm)
 - SAR Baseline 150 m
 - Zoom 15x30 m
- Mini-RF can measure topography
 - Interferometry: 15 m/pix spatial, sub-meter vertical resolution
 - SAR-stereo: 50 m/pix spatial, ~10 m vertical
- Hybrid architecture polarimetric SAR
 - Transmit LCP; Receive *H*- and *V*- linear coherently
 - Four Stokes parameters derived
- Low mass and power requirements
 - 13 kg, 150 W



Mini-RF: A Unique State-of-the-Art Orbiting Radar Observatory



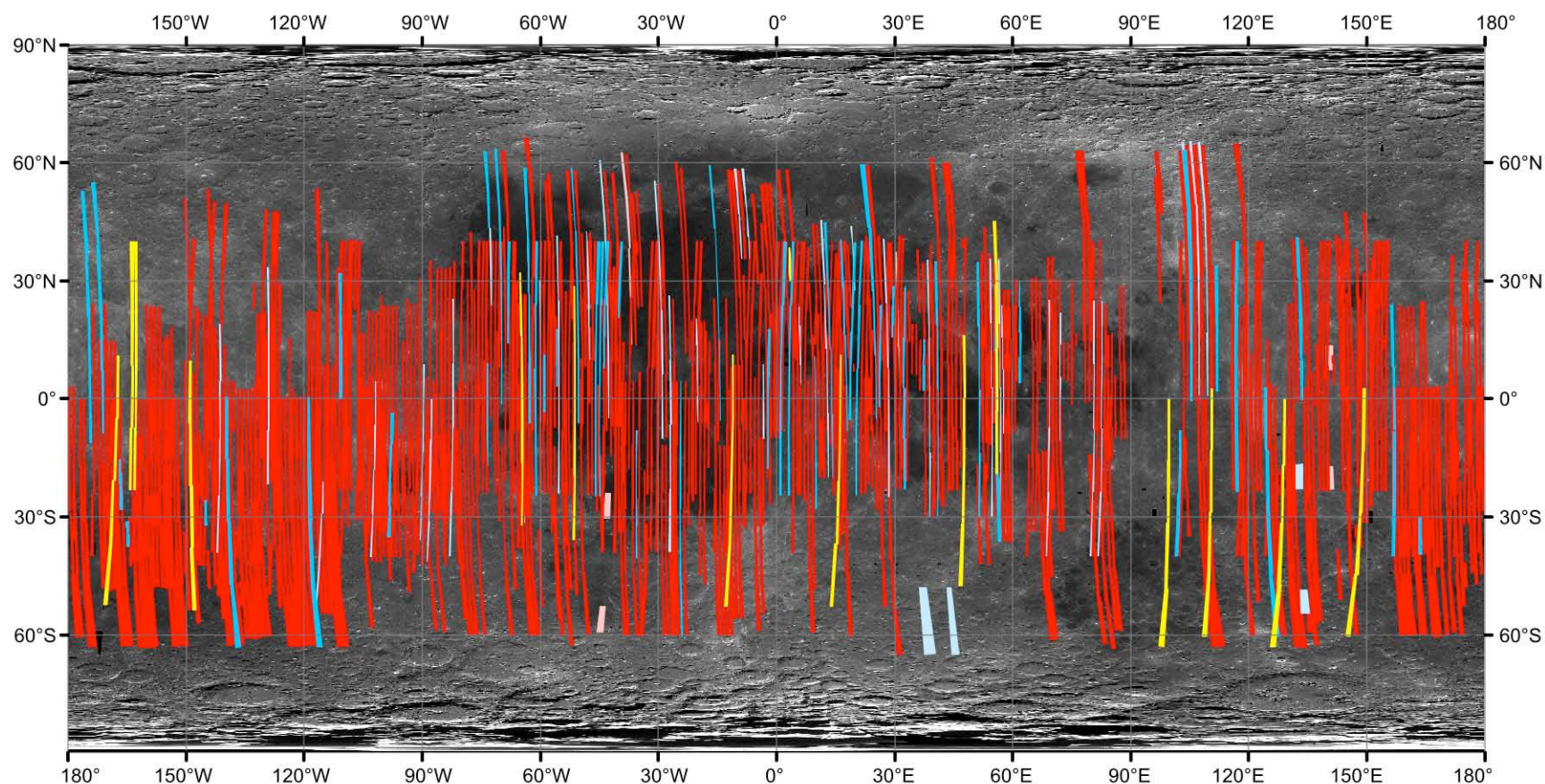
> Mini-RF Imaging Radar on LRO is equivalent (from a science point of view) to the Arecibo-Green Bank Radio Telescope combination

> Mini-RF is the first orbital radar to provide the 4-element Stokes vector at two different wavelengths

> Mini-RF has equal viewing opportunities at the lunar poles, and the entire surface including the far side



Non-Polar Coverage up to May 25 2010

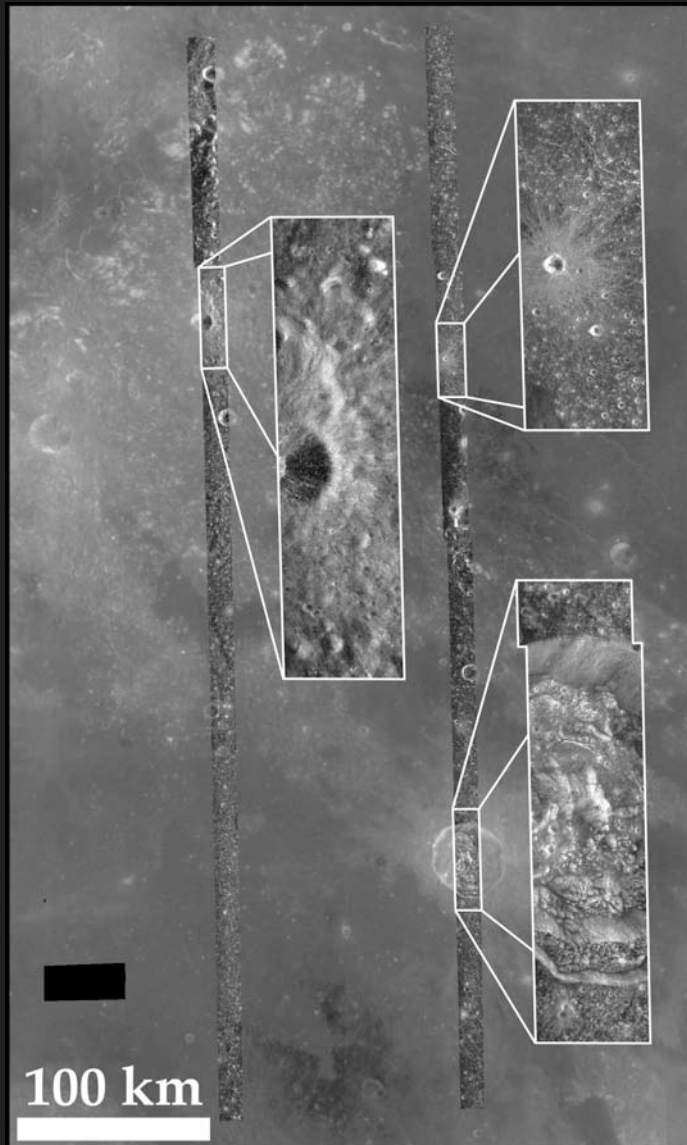


LRO Mini-RF SAR Data Mode

(non-polar data as of 2010-05-20)

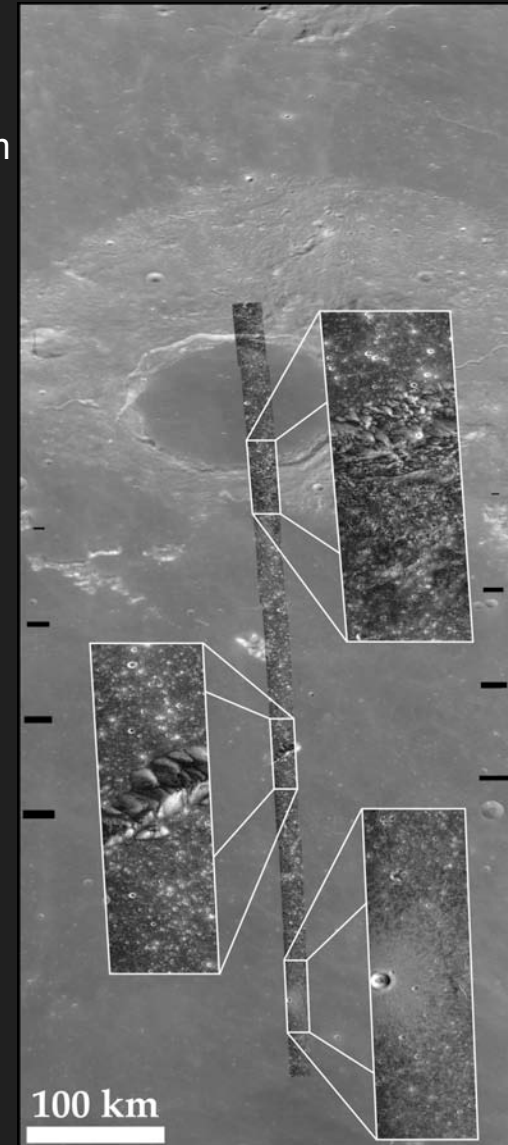
- LRO X Zoom
- LRO X Baseline
- LRO S Zoom
- LRO S Baseline
- gimbal-limited collects

Examples of Mini-RF Data

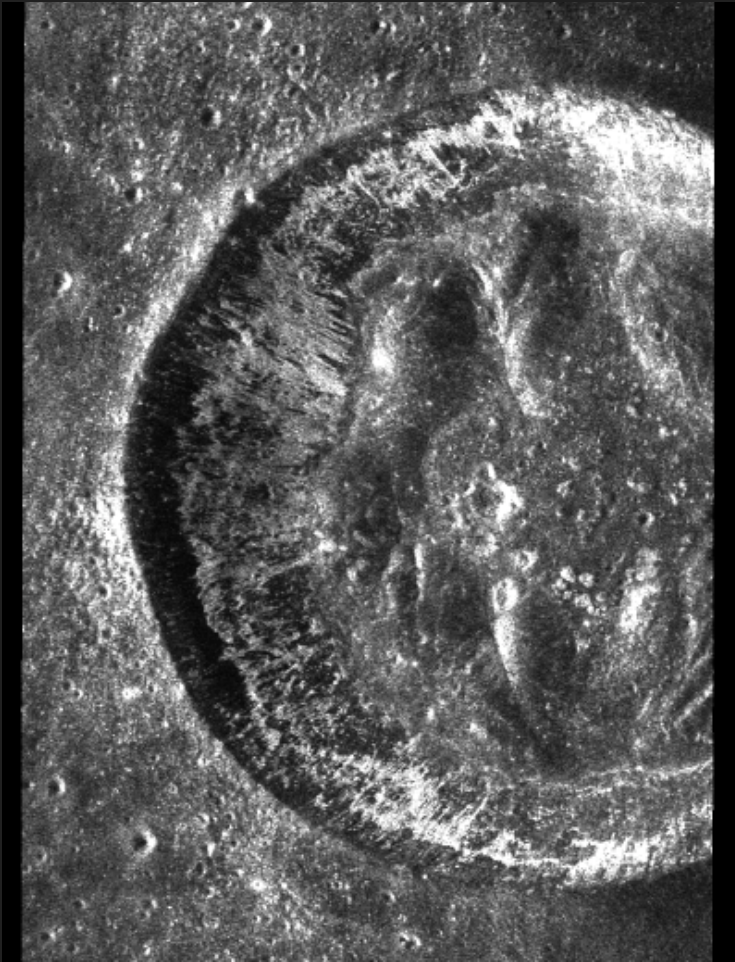


Typical strip is 1920 x 10 km

Mini-RF SAR data co-registered with Clementine UVVIS context mosaics.

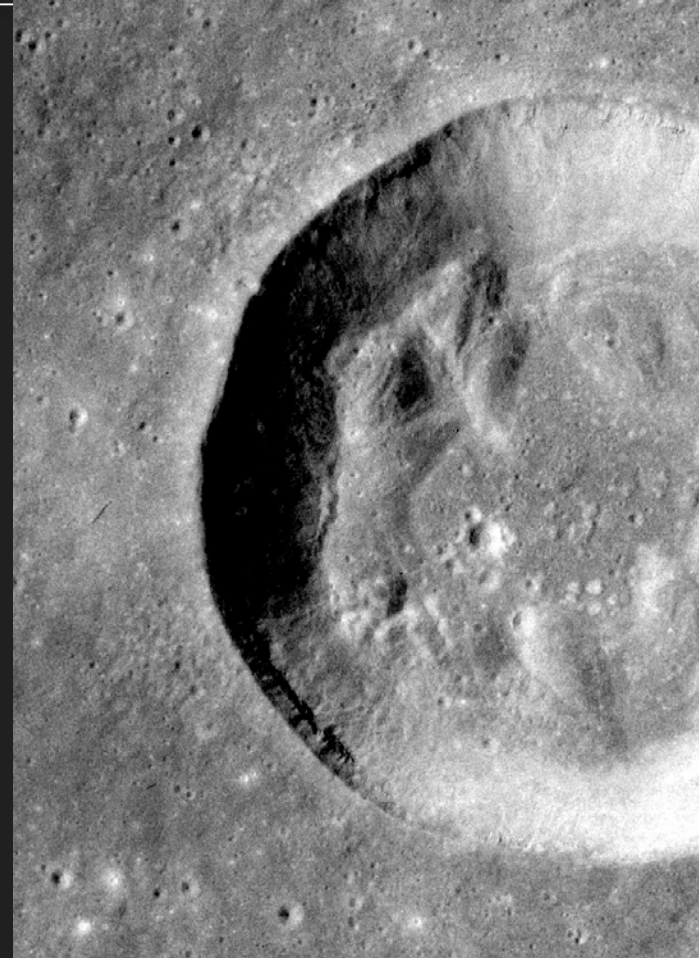


Mini-RF & Optical Data are Synergistic



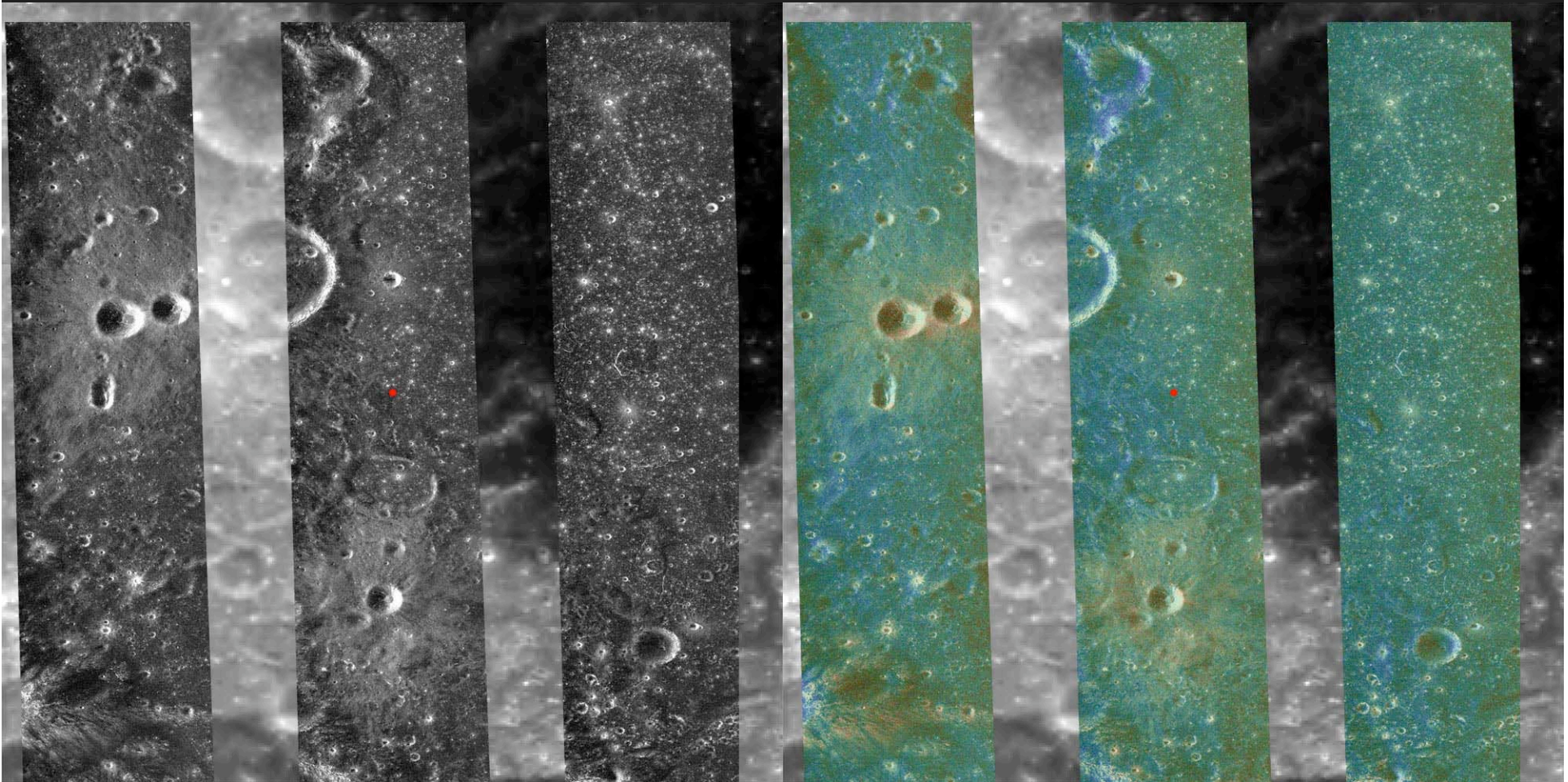
Mini-RF OC S-band zoom

Bessel crater, D=16 km 21.8°N 17.9°E



Apollo 15 Pan frame AS15-1123

Apollo Basin



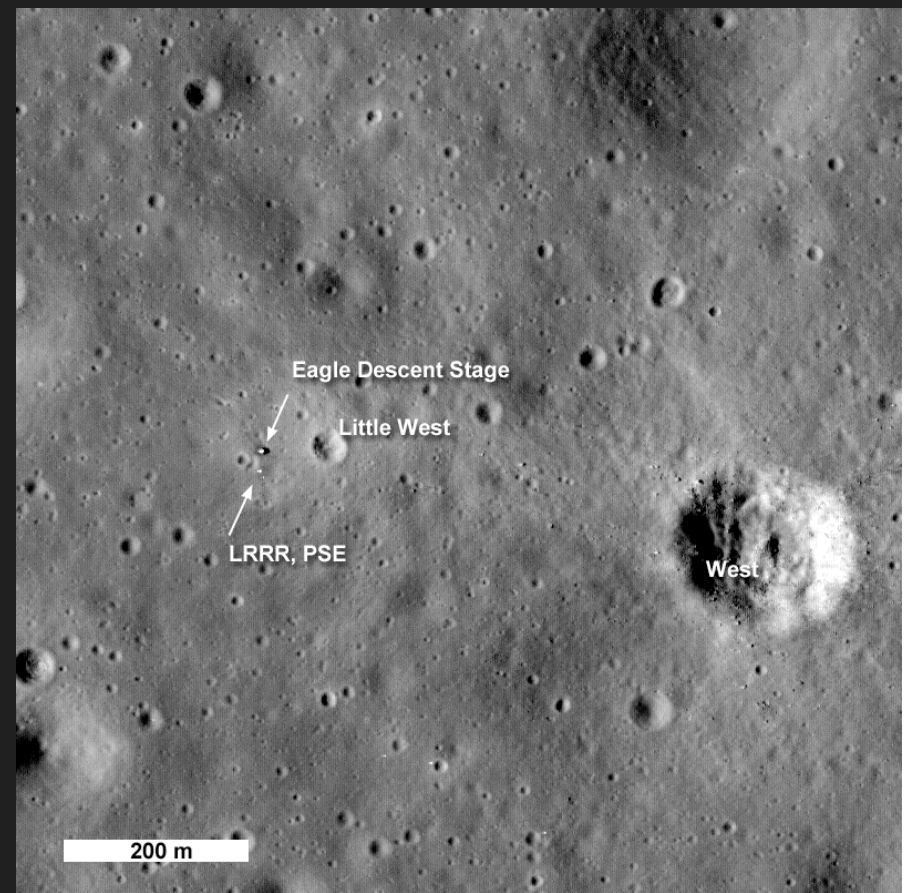
Total radar power over Clementine

Radar CPR over Clementine

Apollo 11

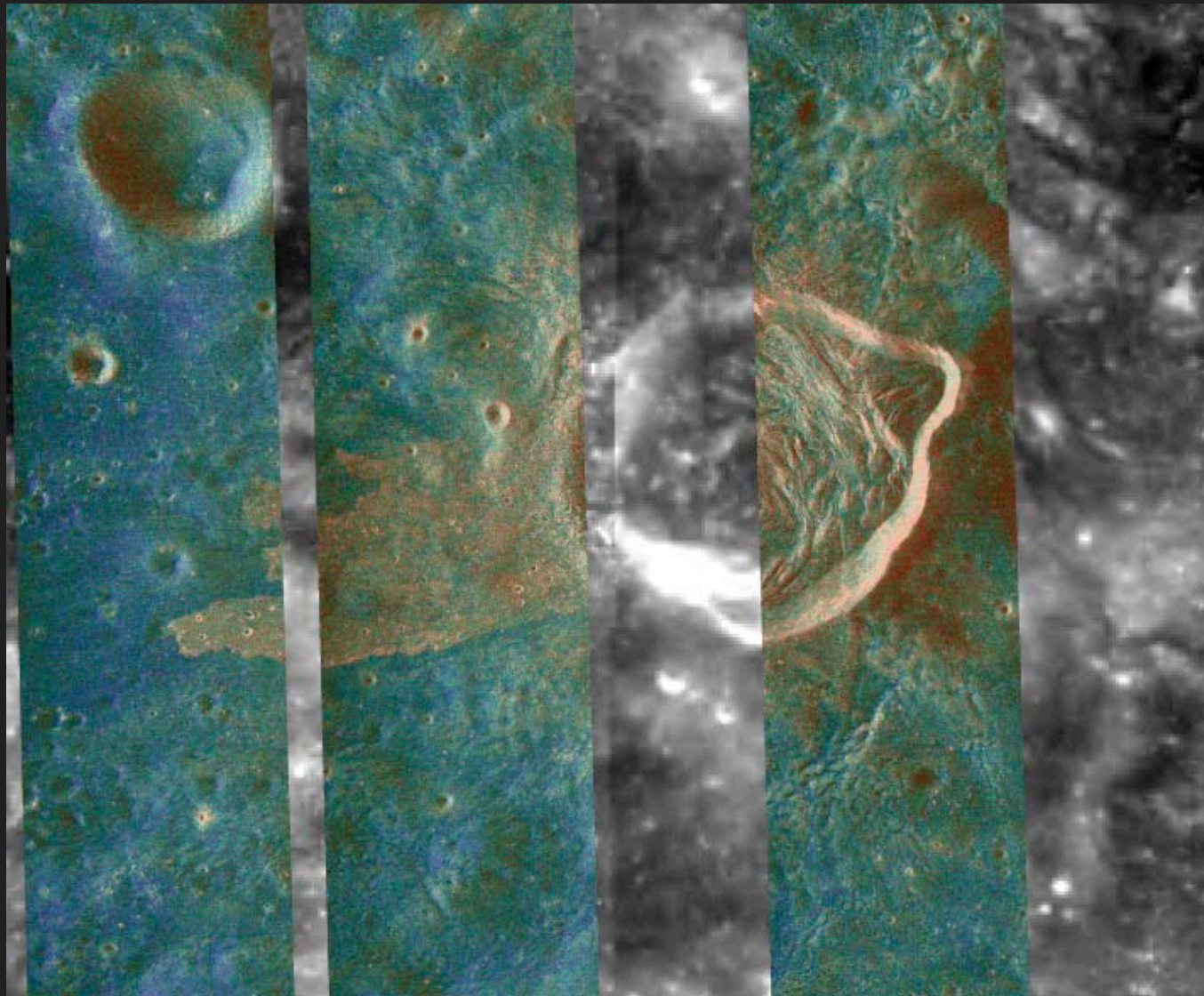


Mini-RF



LROC

Impact Melts



CPR over
total radar
power over
Clementine

Gerasimovich D

Impact Melts

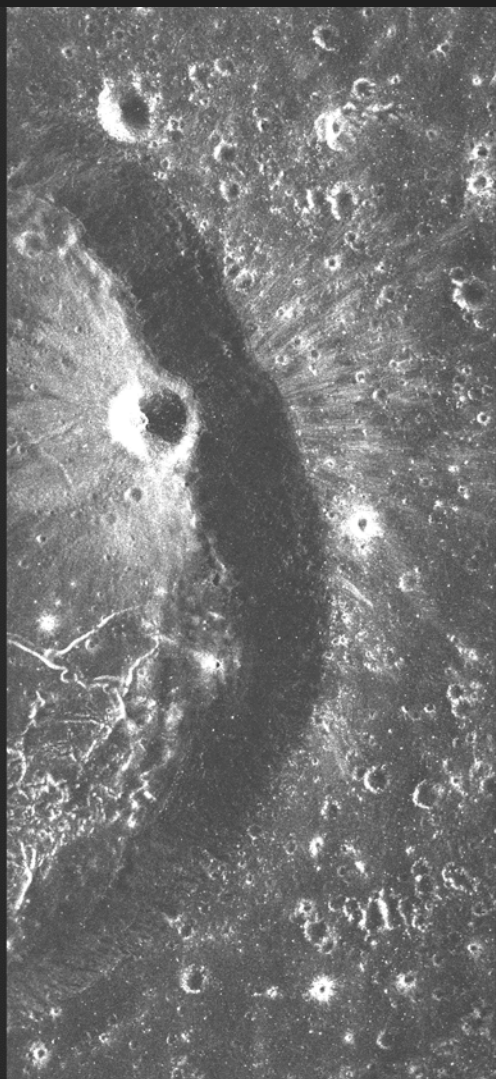


LROC WAC image

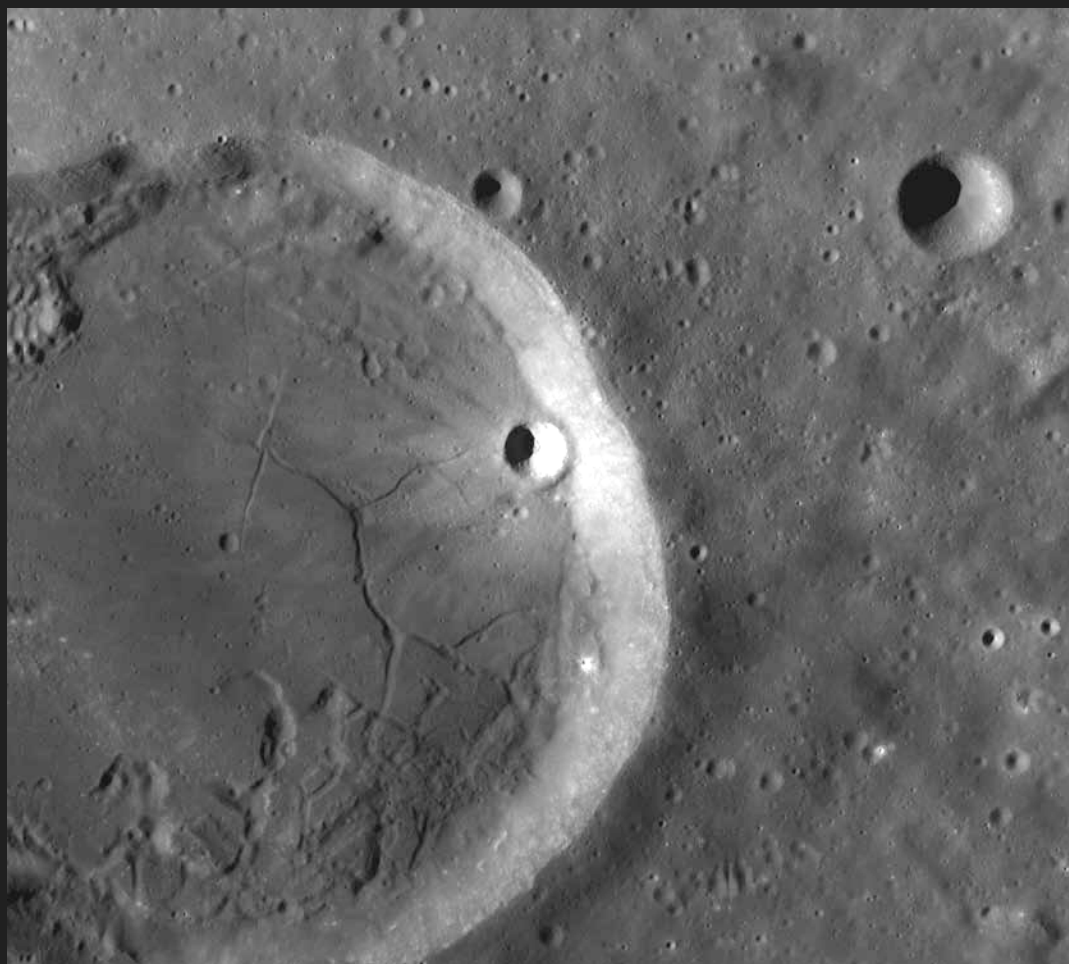
Kopff



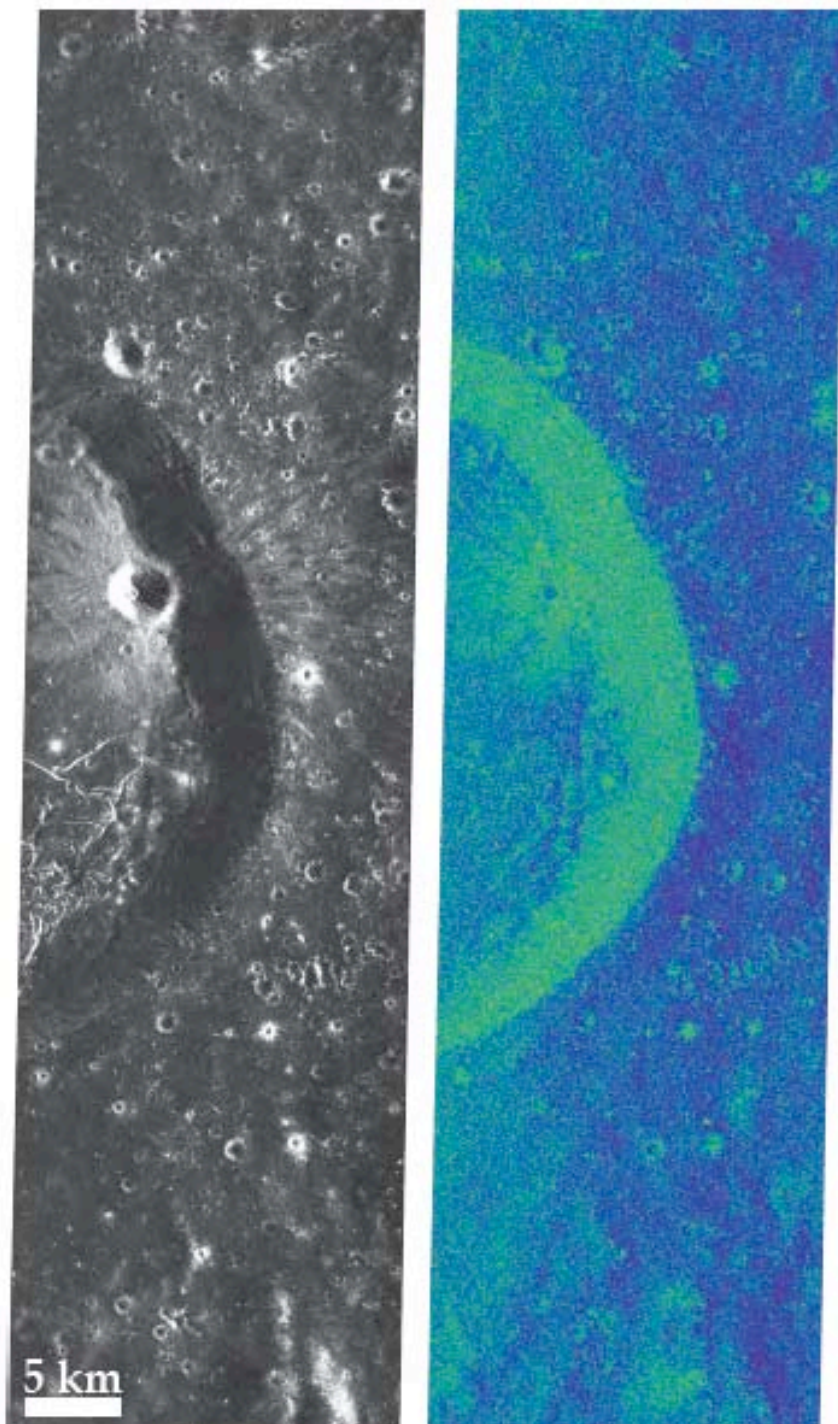
D=16 km 17.4°S 89.6°W



Mini-RF S1



Kaguya TC Mosaic



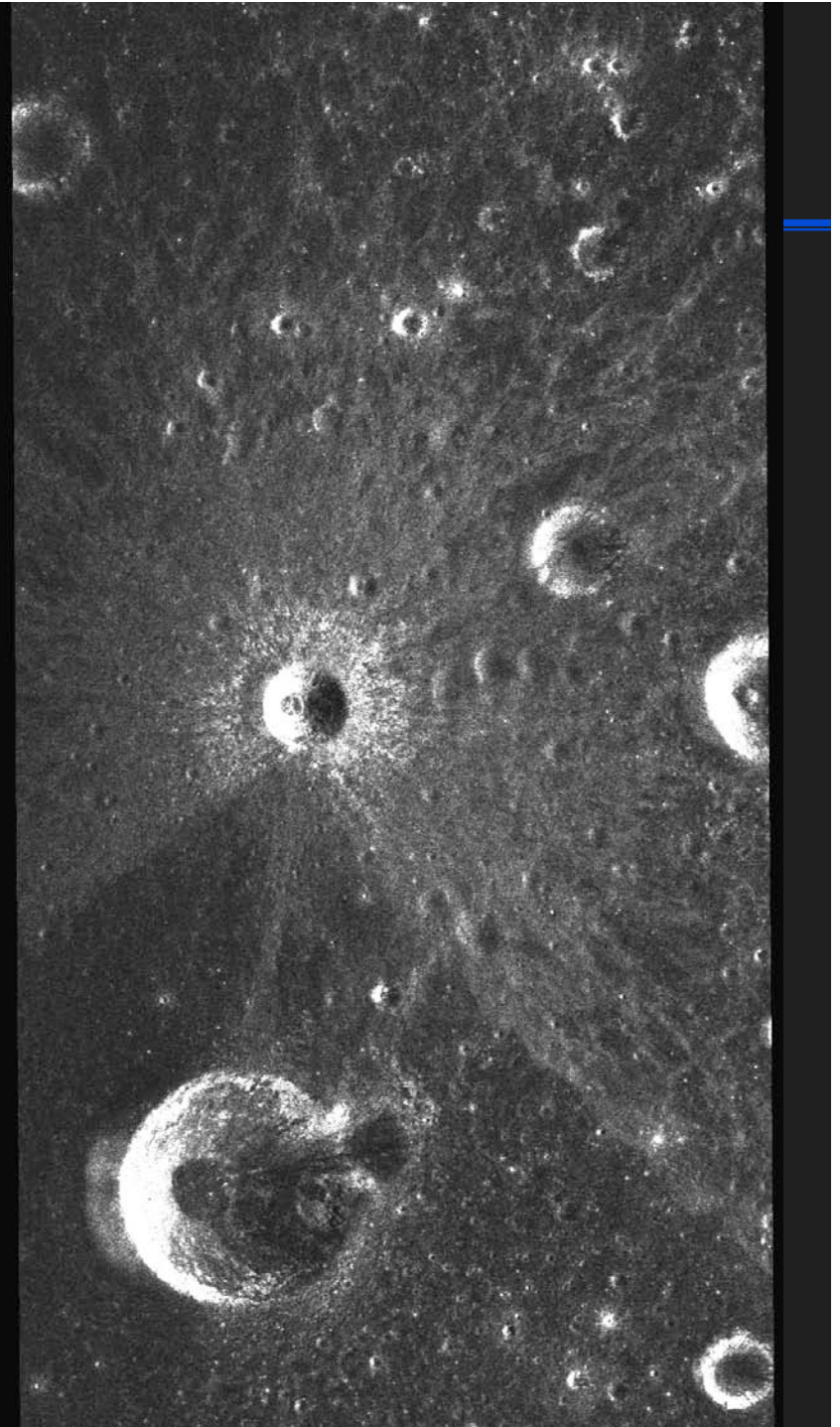
Mini-RF Reveals Information on Impact Process

Kopff crater, D=16 km
17.4°S 89.6°W

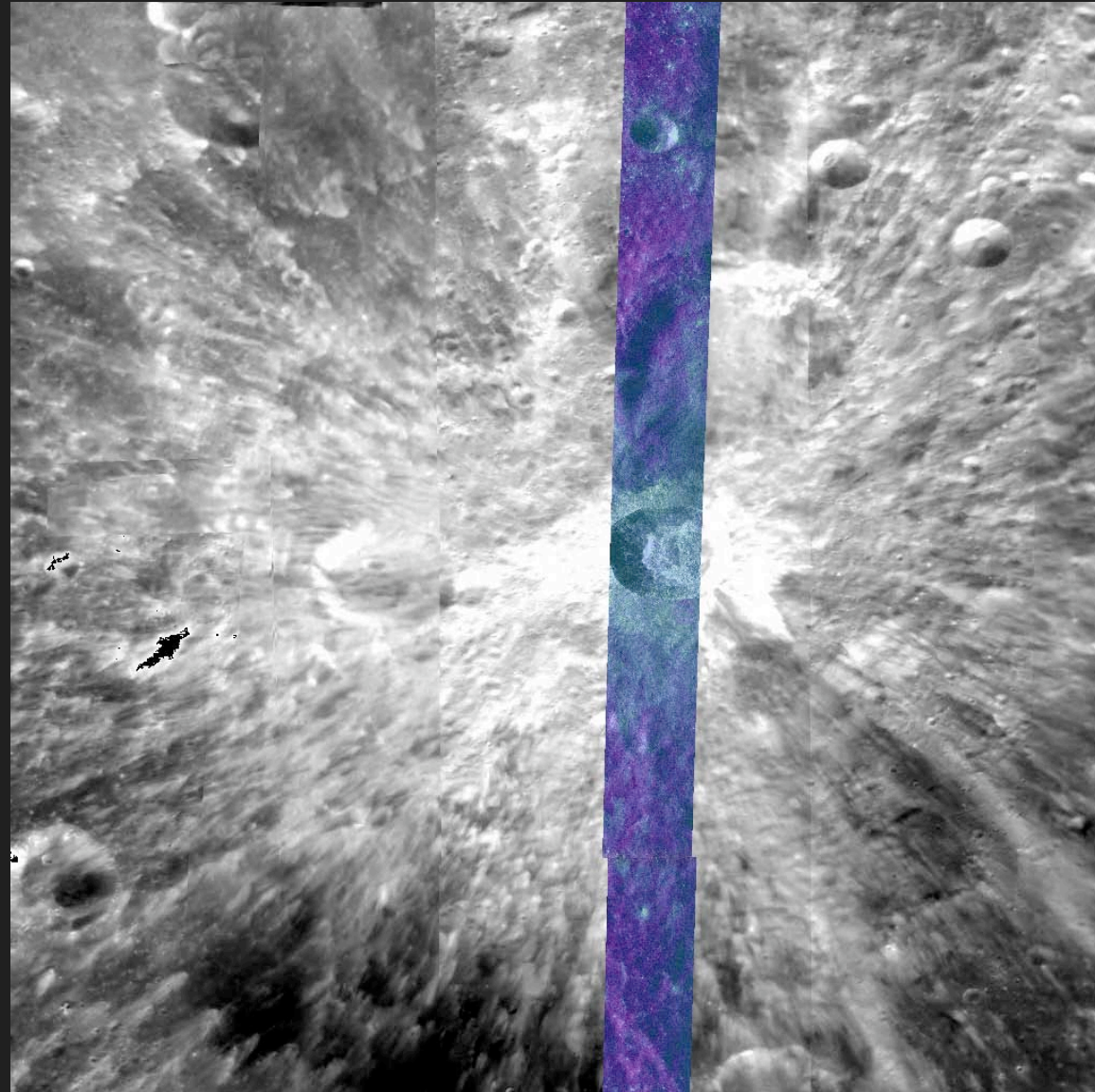
Mini-RF Maps Crater Ejecta Blankets

- Mini-RF maps surface roughness of the lunar surface. In doing so it can map the ejecta blankets that surround impact craters
- Mini-RF can see ejecta that is not visible in optical images.
- This could be because the extended ejecta causes subtle variations in surface roughness or that Mini-RF is detecting subsurface effects.
- Mapping the continuous and discontinuous ejecta blankets that surround impact craters will help us to better understand the physics of impact cratering.

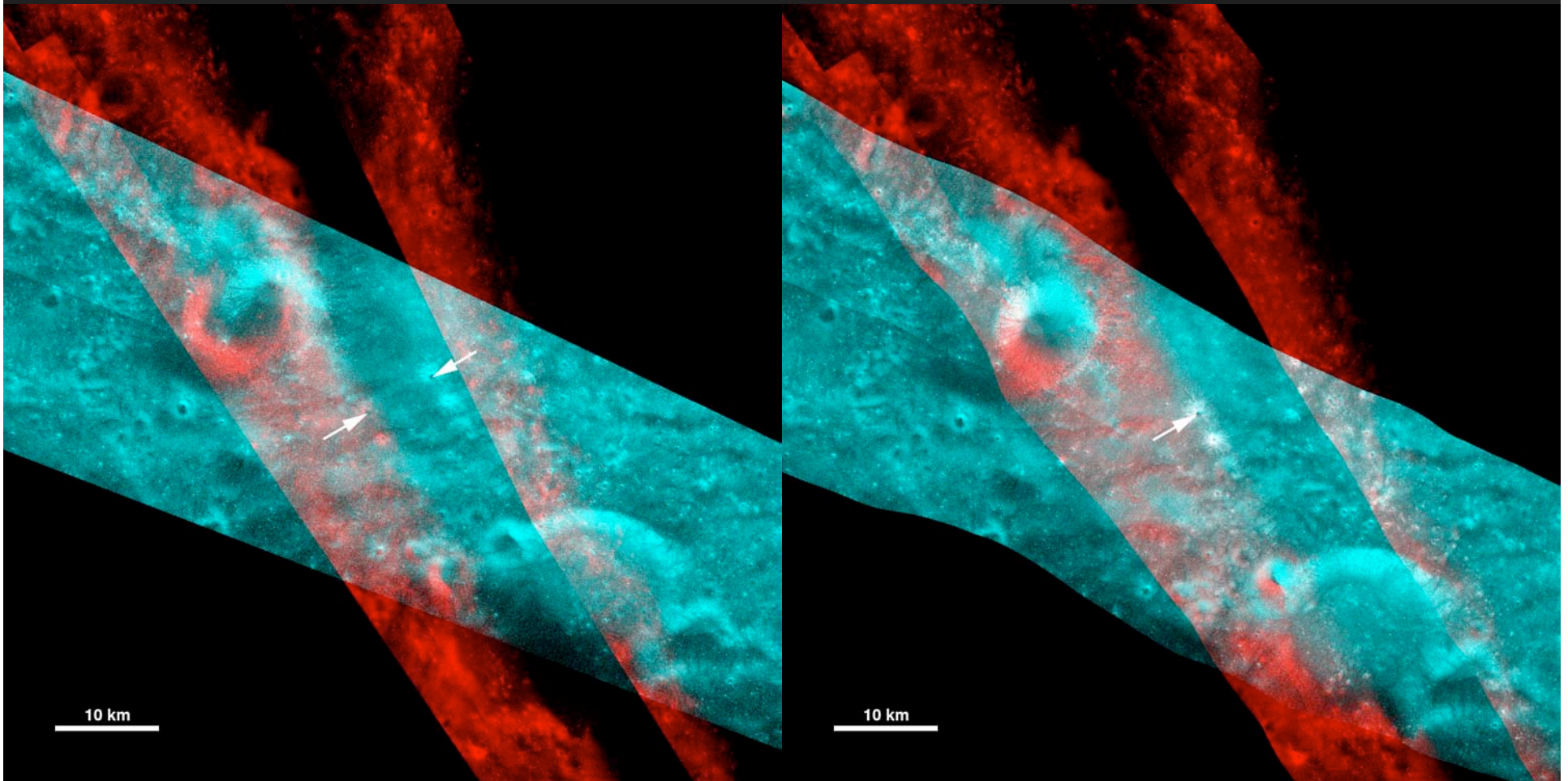
Oblique impact crater in mare Nubium. Both the continuous and extended ejecta blankets are detected



Giordano Bruno



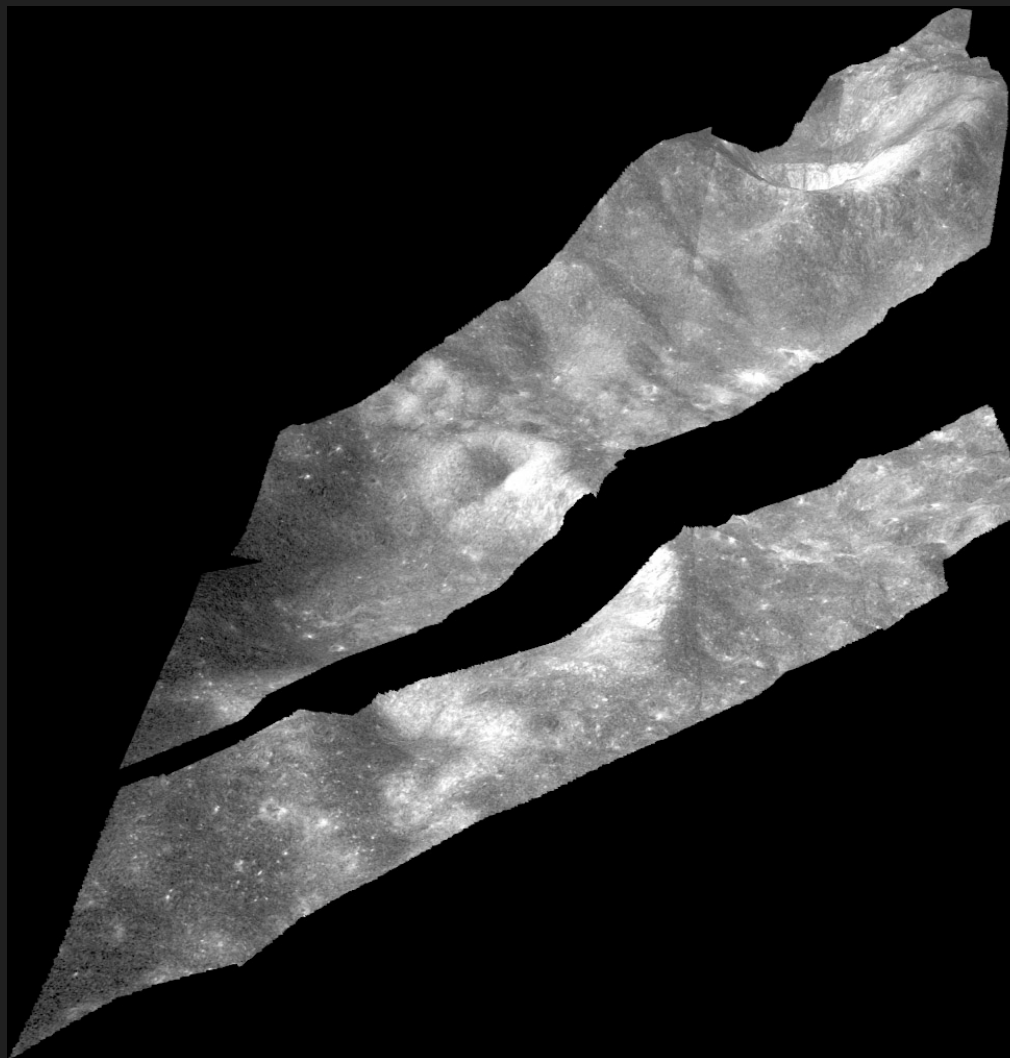
Uncontrolled v. Controlled



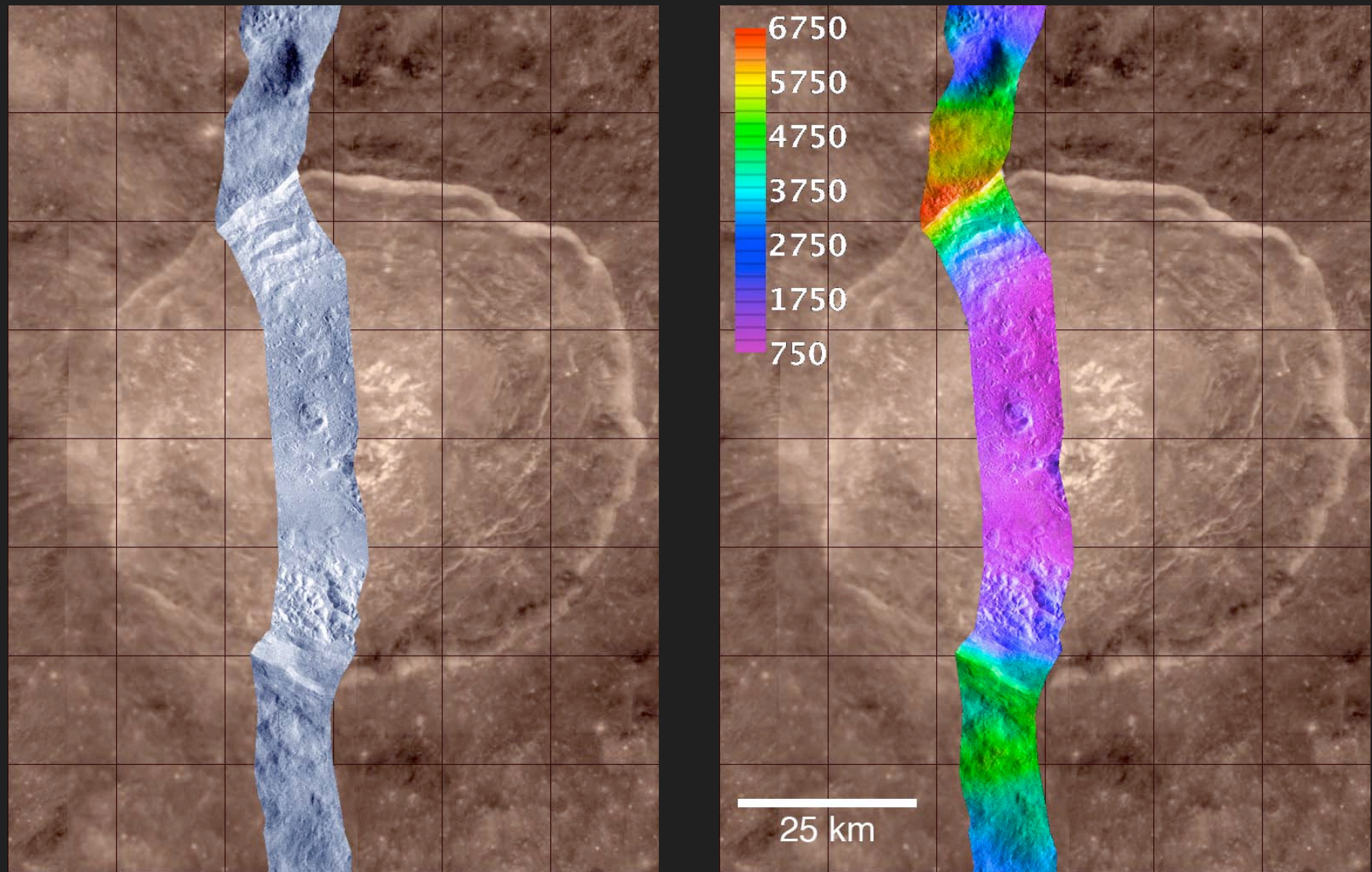
Uncontrolled mosaic, projected onto sphere. Offset between 2 sets of orbits is nearly 10 km

Controlled mosaic, projected onto Kaguya altimetry (ortho-rectified) reduces mismatch to too small to see

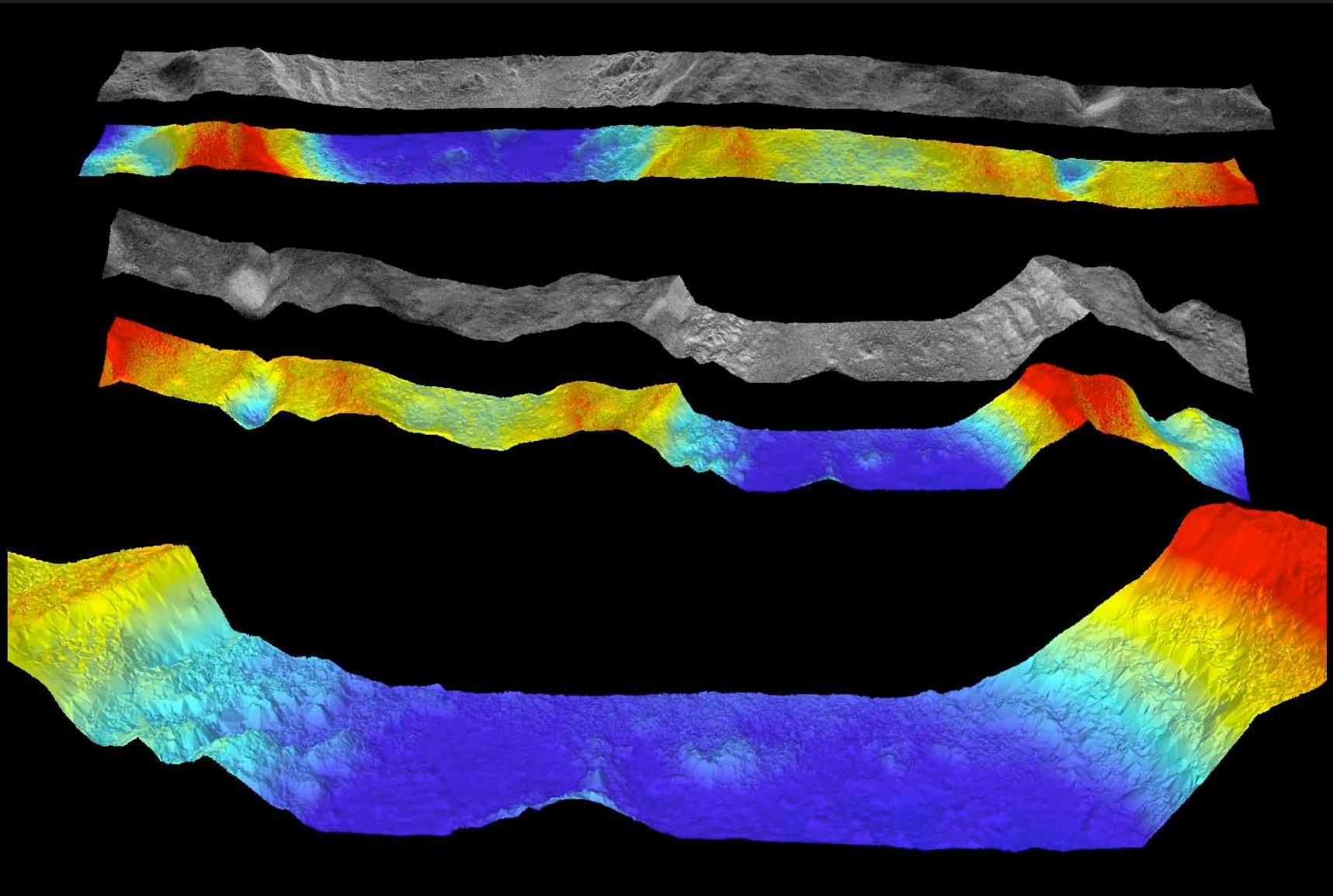
Cabeus Perspective View with Orthoimage Mosaic



Roll-Derived Stereo of Jackson Crater



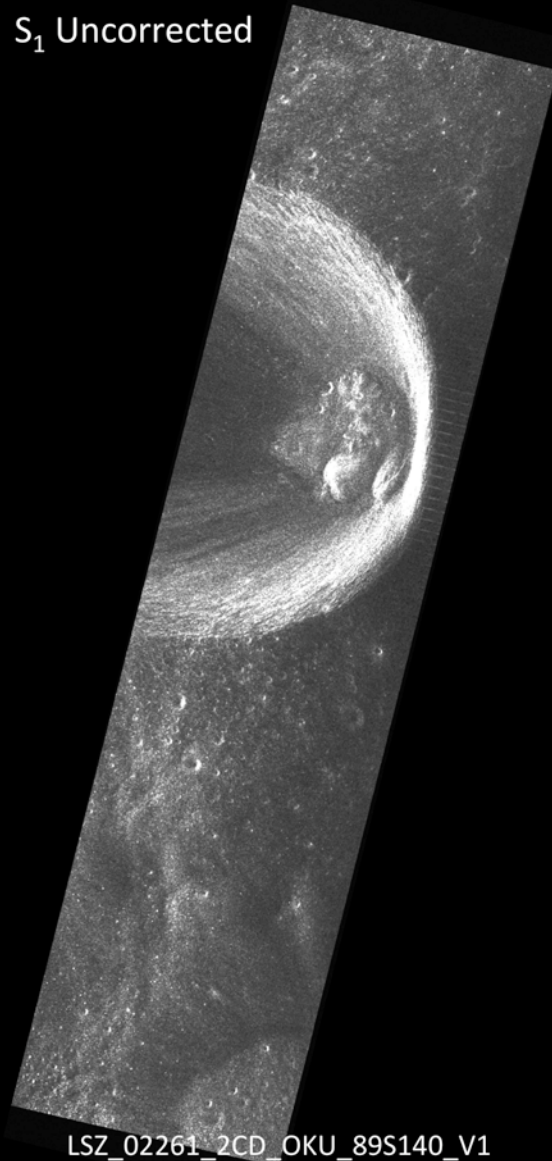
Roll-Derived Stereo of Jackson Crater



Shackleton

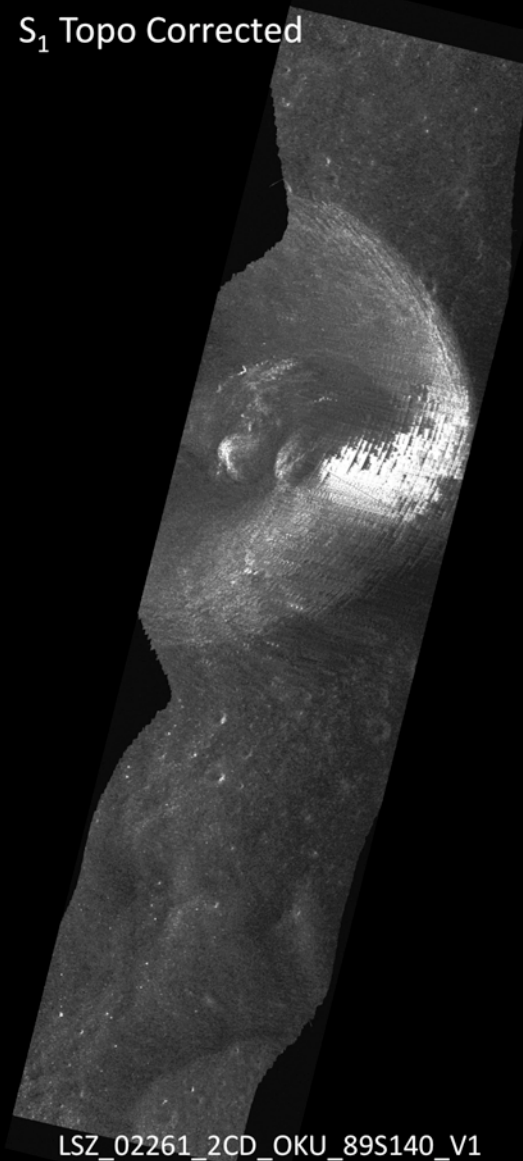


S₁ Uncorrected



LSZ_02261_2CD_OKU_89S140_V1

S₁ Topo Corrected



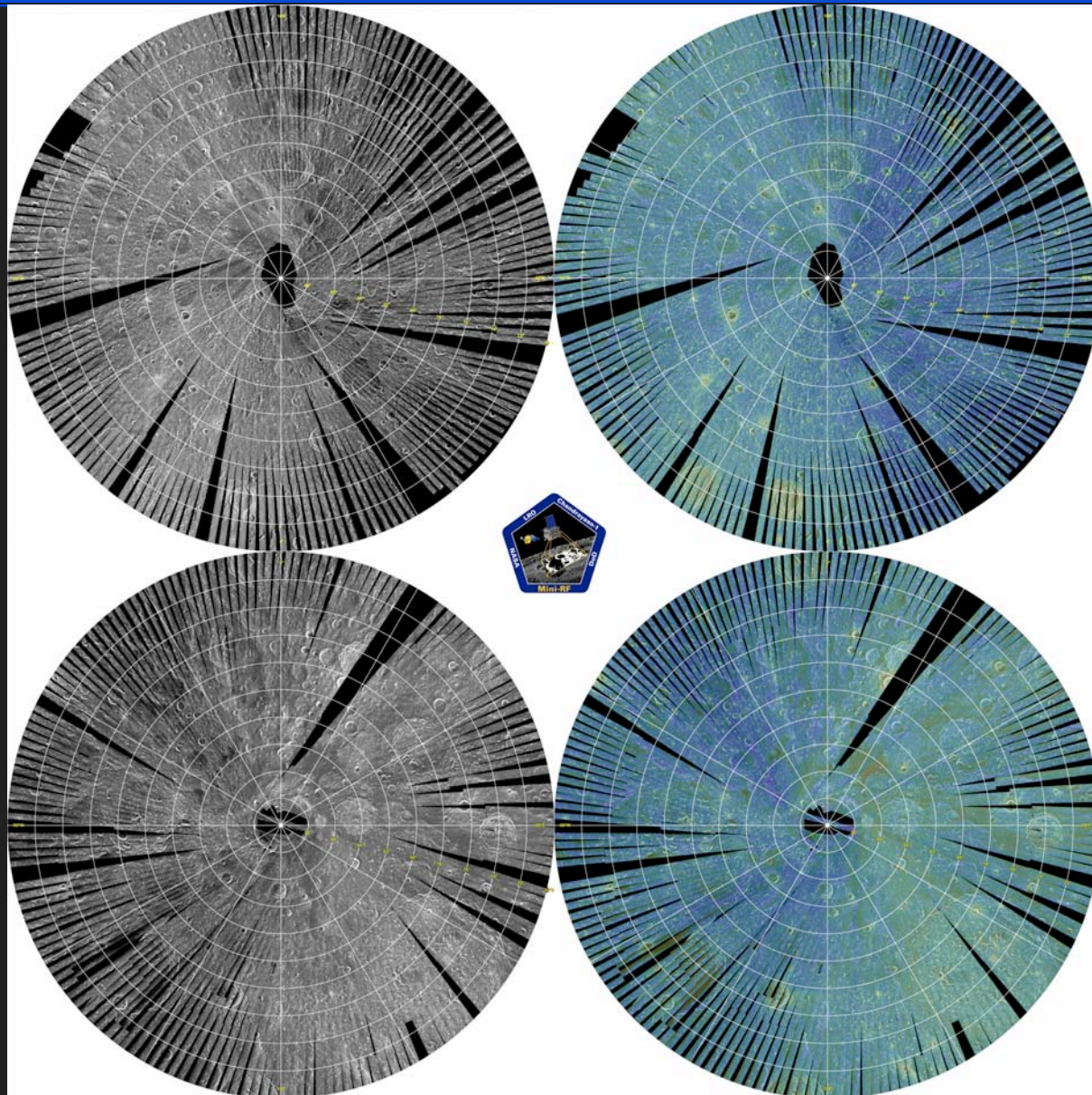
LSZ_02261_2CD_OKU_89S140_V1

Polar Campaign

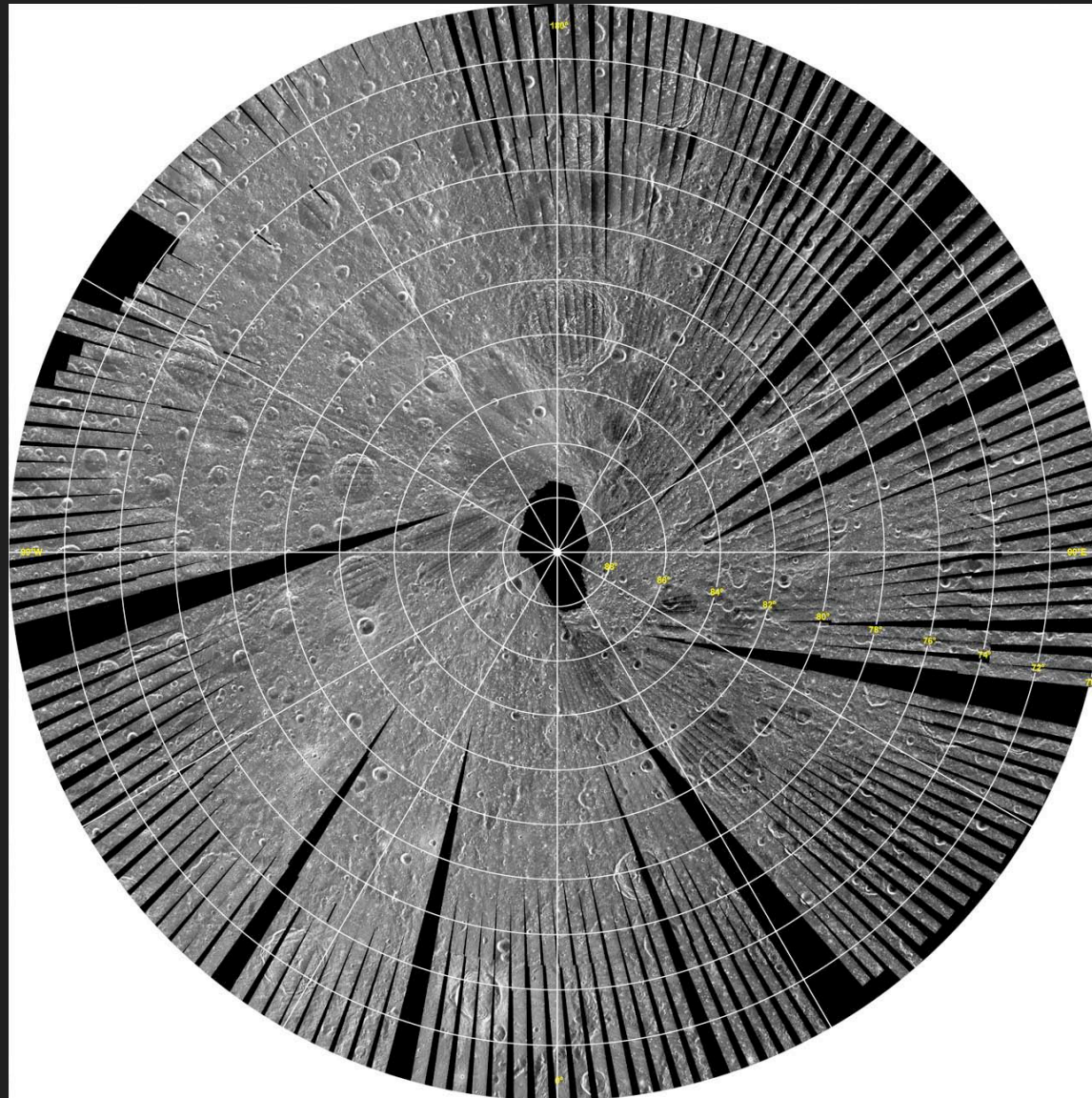


- During times of high solar beta ($> 60^\circ$) the solar array remains in a locked position that is congruent with Mini-RF operations.
- Mini-RF is currently a little over half way through its first high-beta polar-mapping campaign
- We have been concentrating on acquiring S-zoom coverage

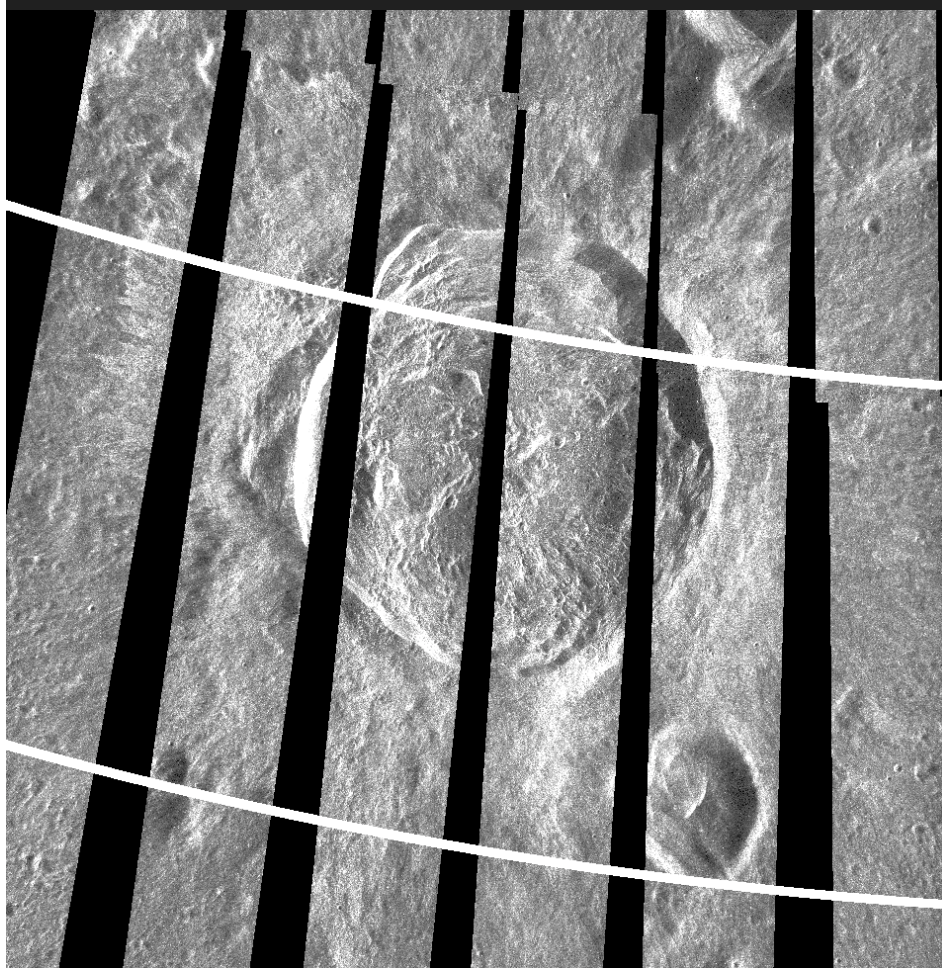
Polar Campaign



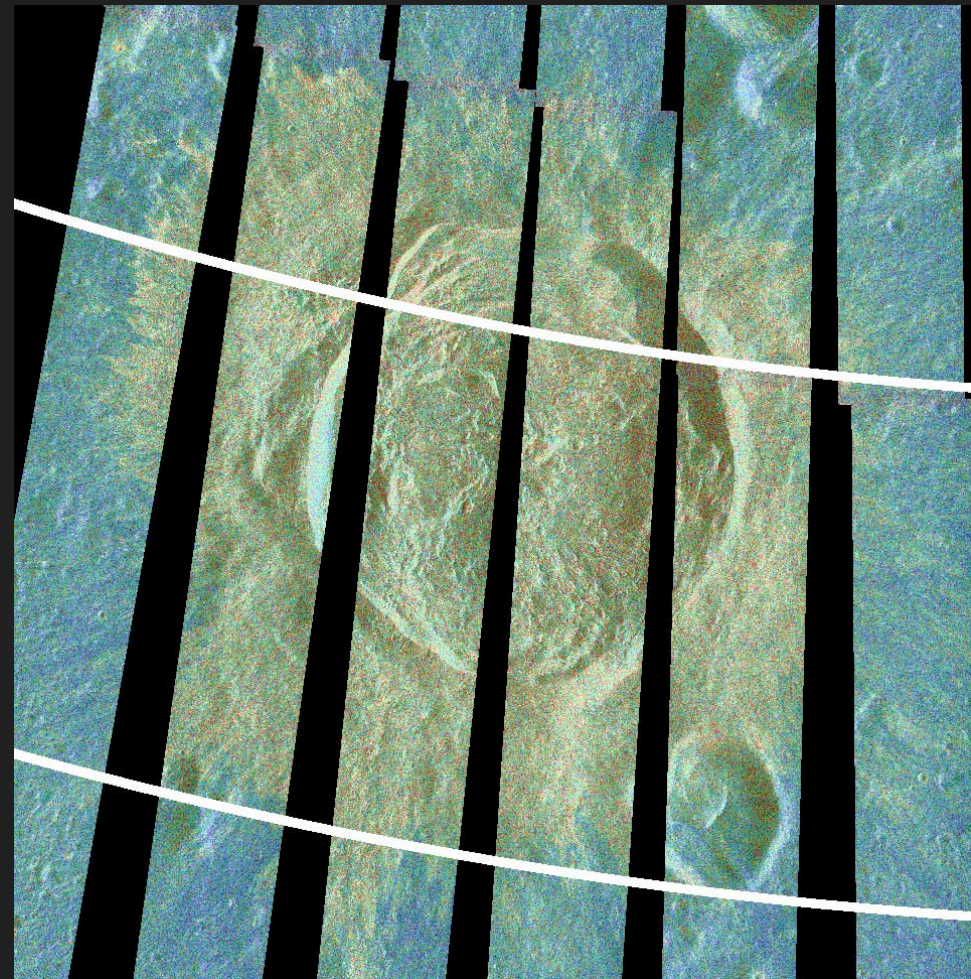
North Pole



Anaxagoras



S1



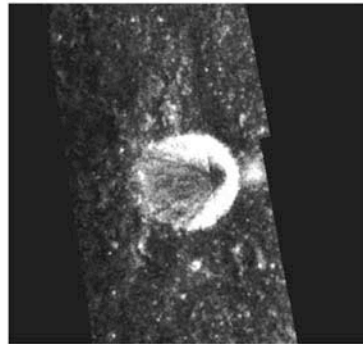
CPR over S1

Rozhdestvensky N

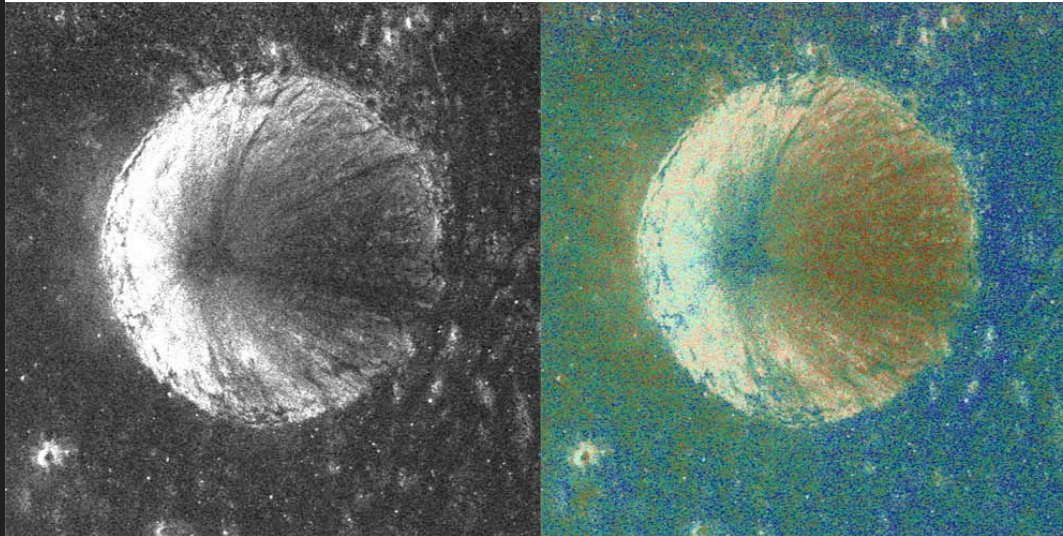
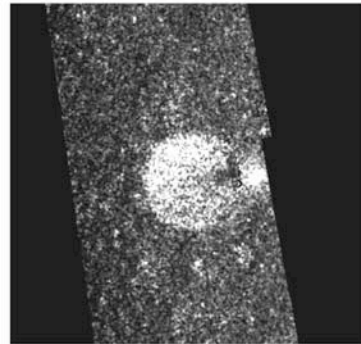


Anomalous crater on floor of Rozhdestvensky (polar)

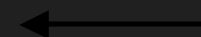
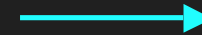
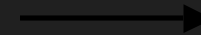
SC



CPR



Illumination direction



Summary



- Mini-RF is a highly capable instrument, obtaining data about lunar surface properties, nature of polar volatiles, surface topography, etc.
- Data acquisition strategy:
 - Non-polar data continuing at much reduced pace until beta angle $\geq 60^\circ$
 - Polar campaigns to be started at beta angle $\geq 60^\circ$
- PDS Delivery successful, >15TB of Mini-RF sent to PDS
- Please use the data!

Tessera

